

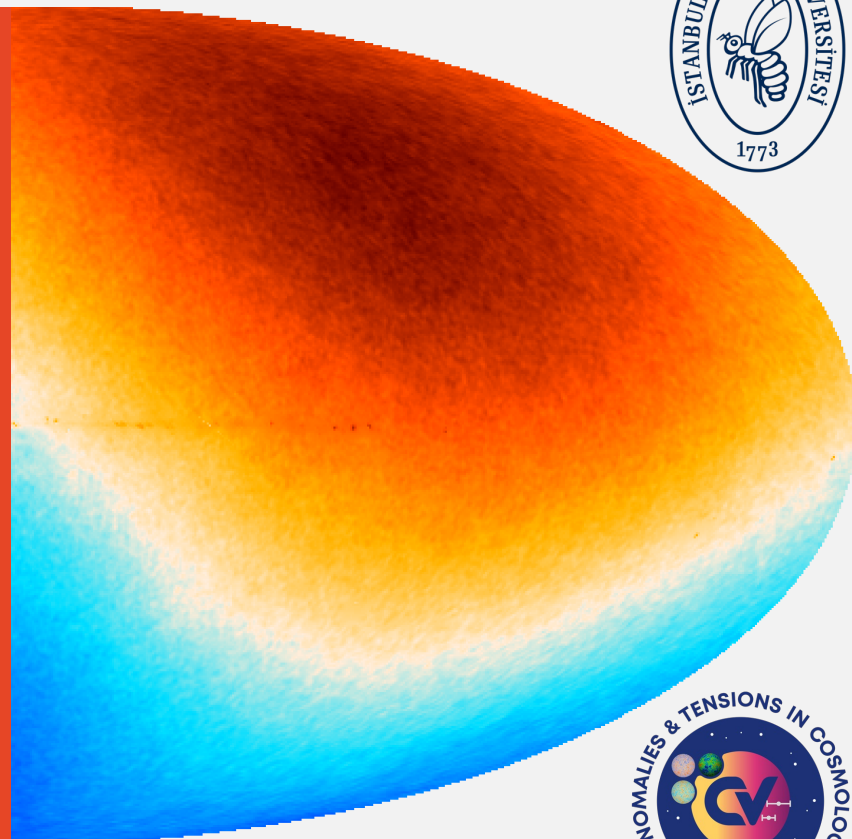
# Cosmic Dipole Tensions: Confronting Planck, NVSS, RACS and CatWISE

*CosmoVerse@Istanbul 2025*

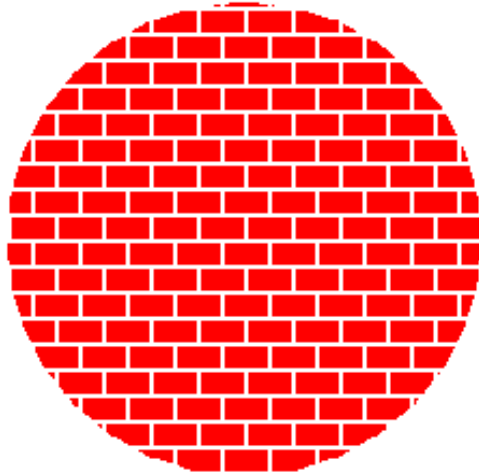
**Mali Land-Strykowski** | Sydney Institute for Astronomy  
Supervised by Geraint F. Lewis and Tara Murphy



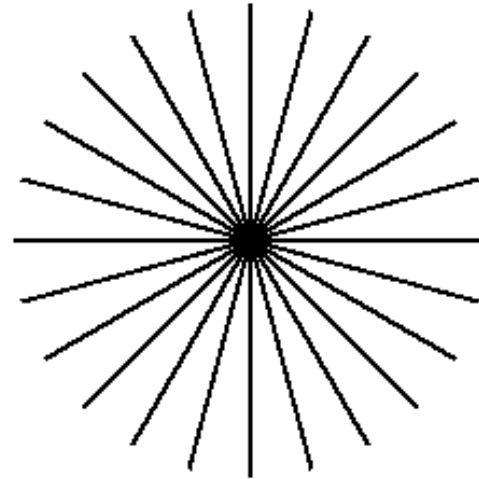
*Celebrating 175 years*



# The Cosmological Principle: a big assumption



Universe is **homogeneous**  
(same everywhere)



Universe is **isotropic**  
(same in all directions)

Image credit: Edward L. Wright

# The Cosmological Principle: important in $\Lambda$ CDM

- Assumed in the Friedmann–Lemaître–Robertson–Walker spacetime
- Therefore, assumed in the standard concordance model of cosmology,  $\Lambda$ CDM



# The Cosmological Principle: important in $\Lambda$ CDM

- Assumed in the Friedmann–Lemaître–Robertson–Walker spacetime
- Therefore, assumed in the standard concordance model of cosmology,  $\Lambda$ CDM
- If the Cosmological Principle **breaks**, all the stones come **tumbling to the ground**...



We must continue to **interrogate the CP** and its **evidence**.



# The Cosmic Microwave Background: very smooth map (CMB)

$$\Delta T/T \approx 10^{-5}$$

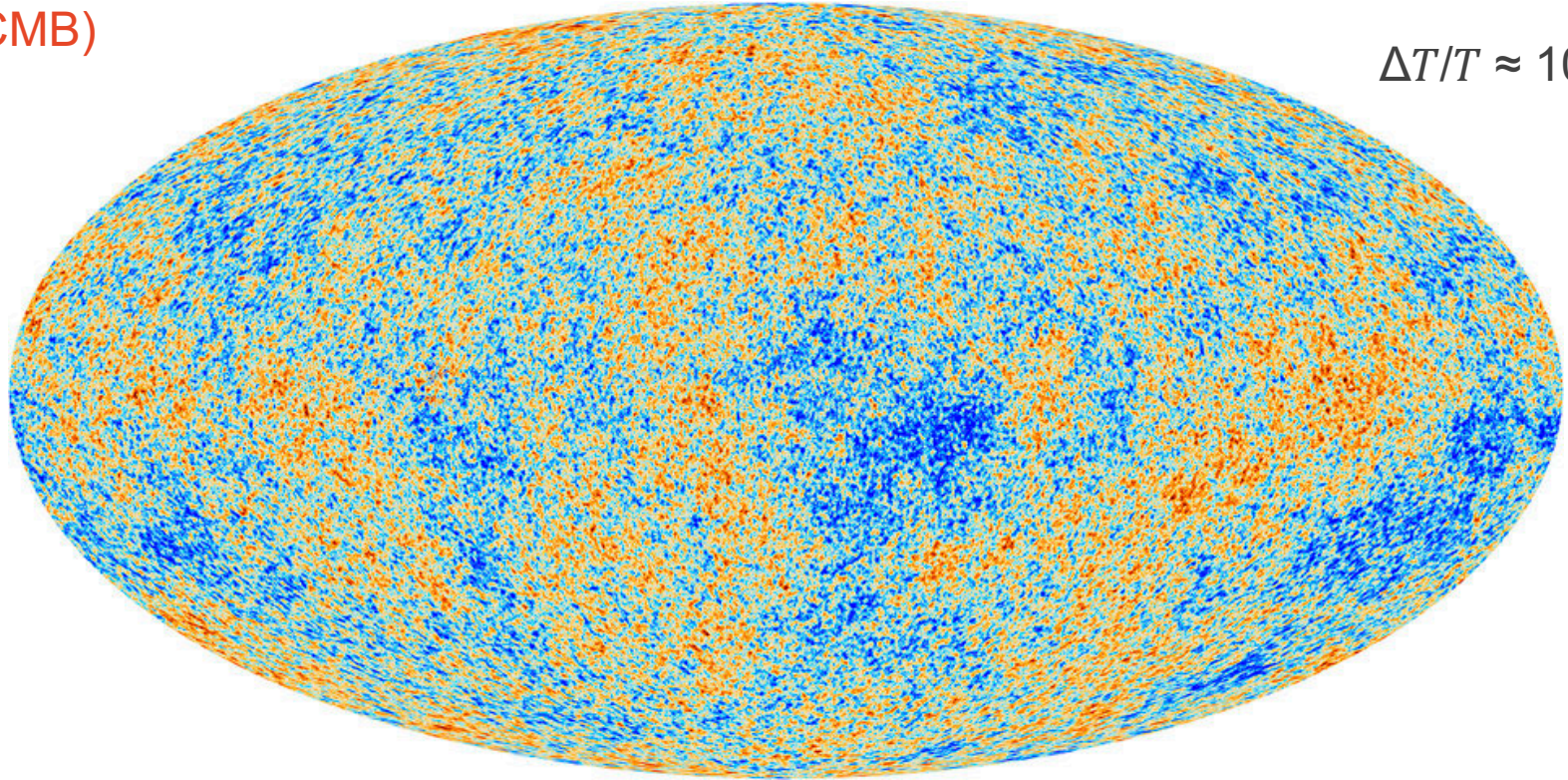
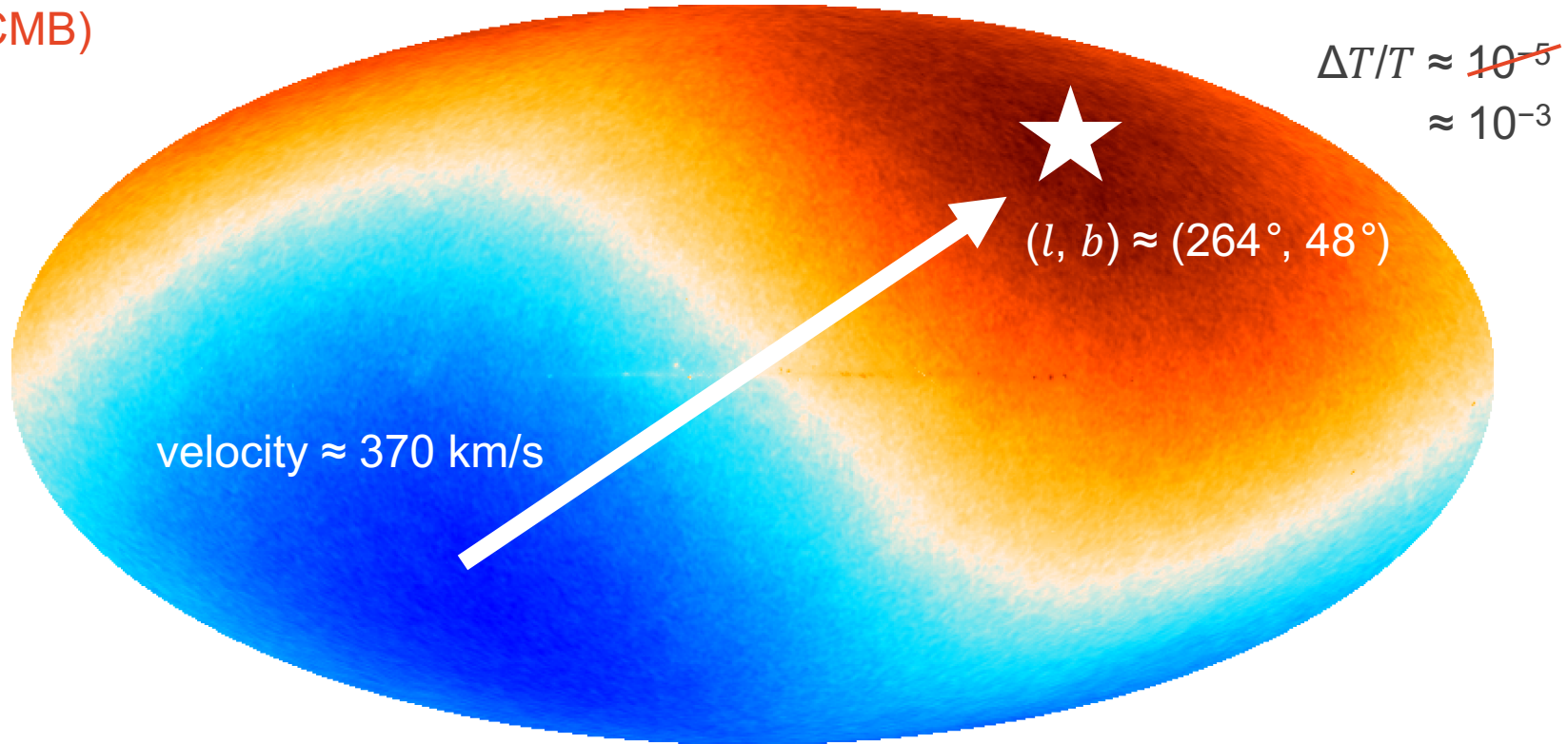


Image credit: *Planck*

# The Cosmic Microwave Background: temperature **dipole** (CMB)



# Relativistic aberration

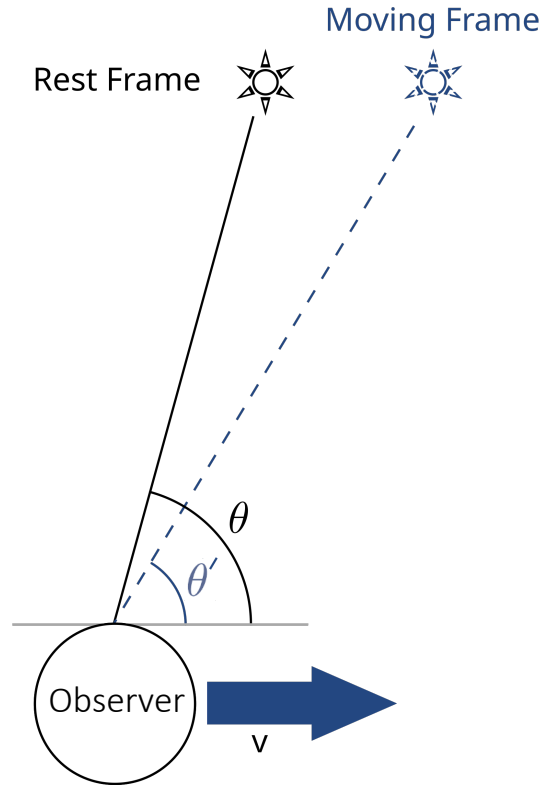
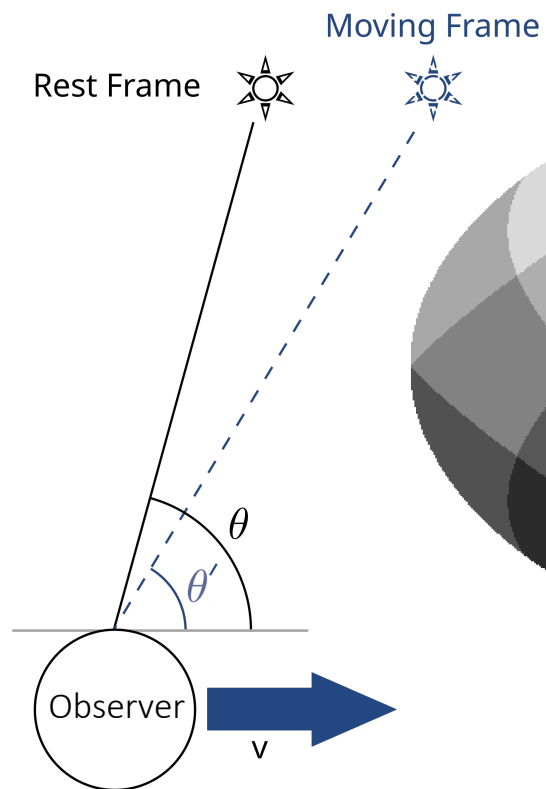


Image credit: Wikipedia

# Relativistic aberration



Sky divided into **equal** angular-area pixels

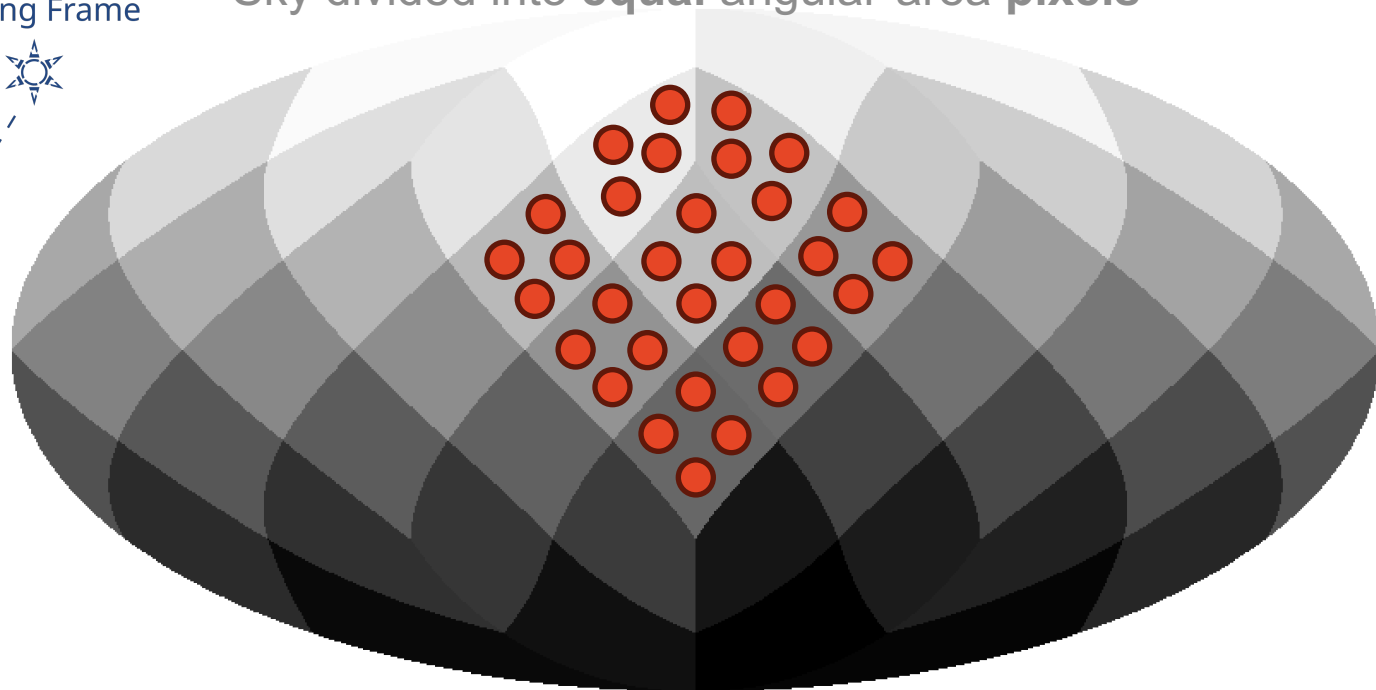
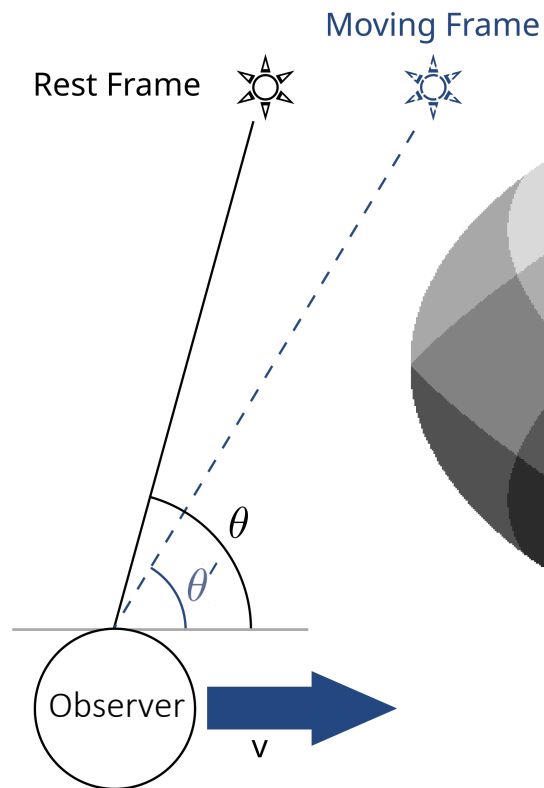


Image credit: Wikipedia

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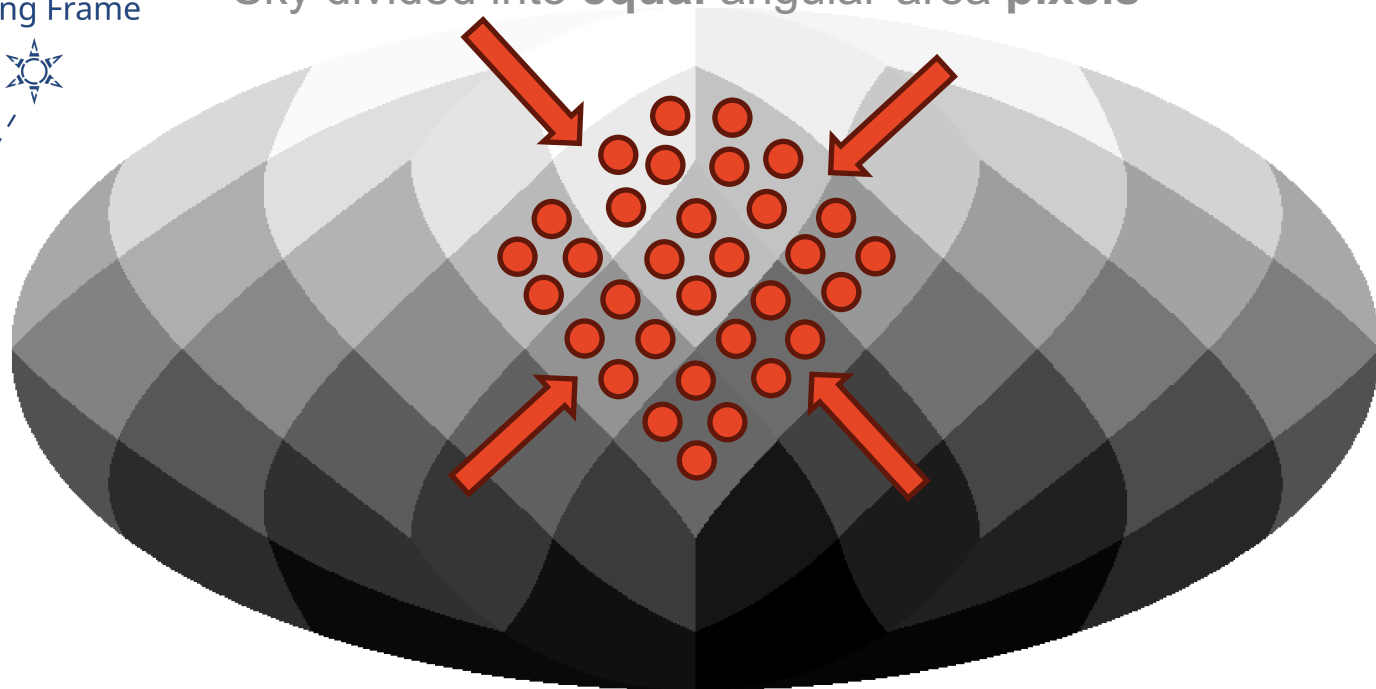
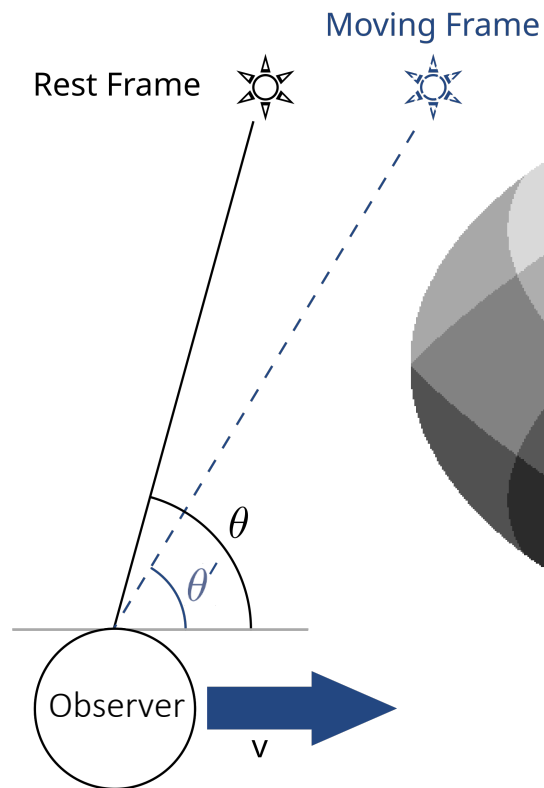


Image credit: Wikipedia



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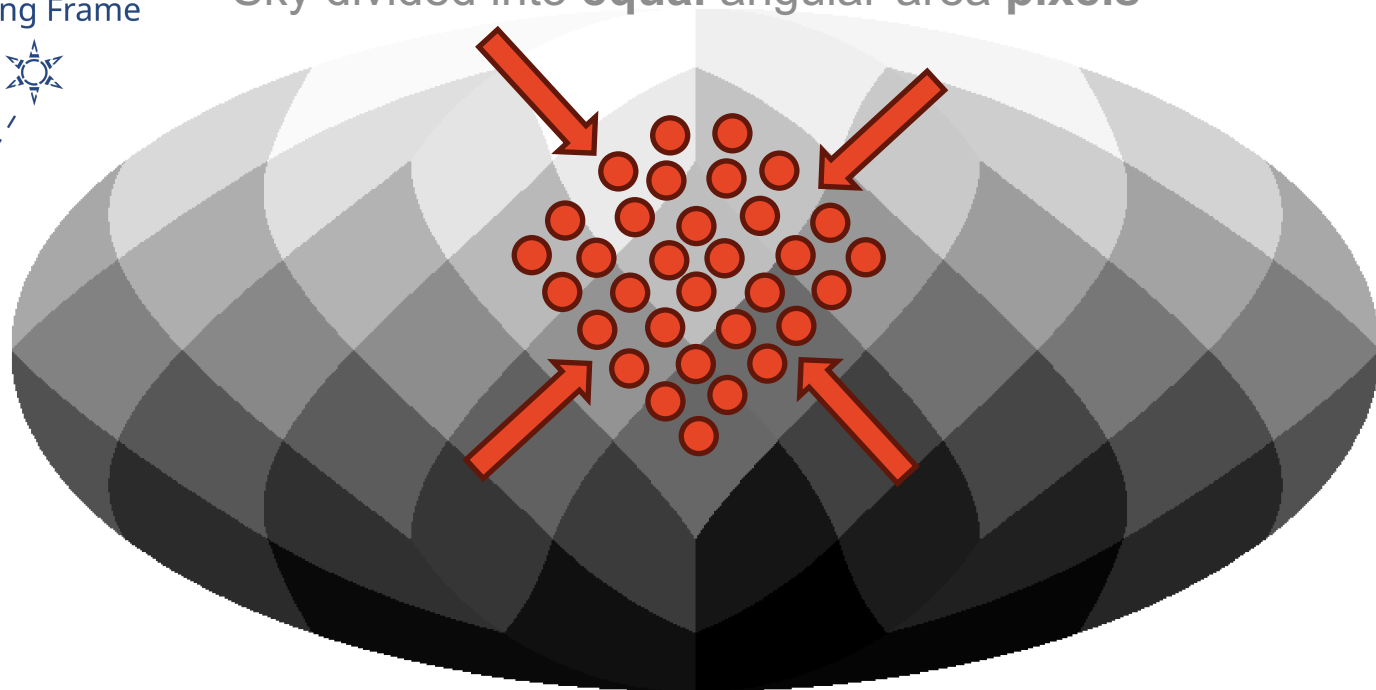
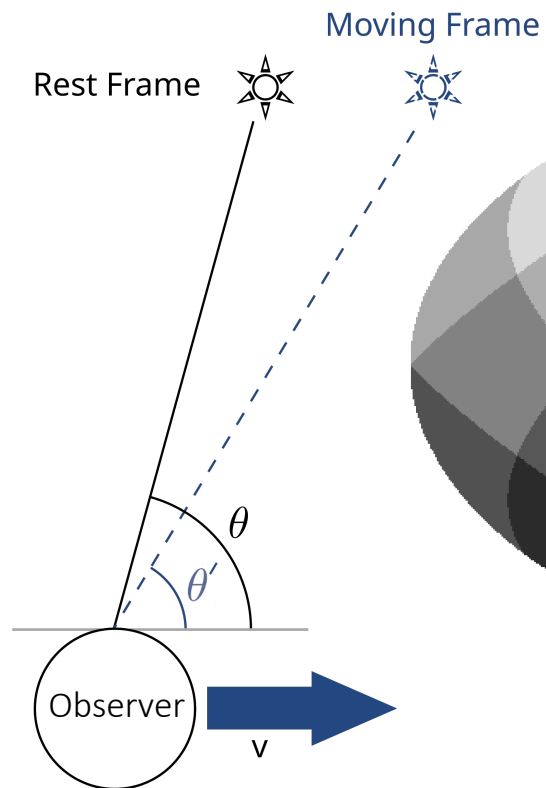
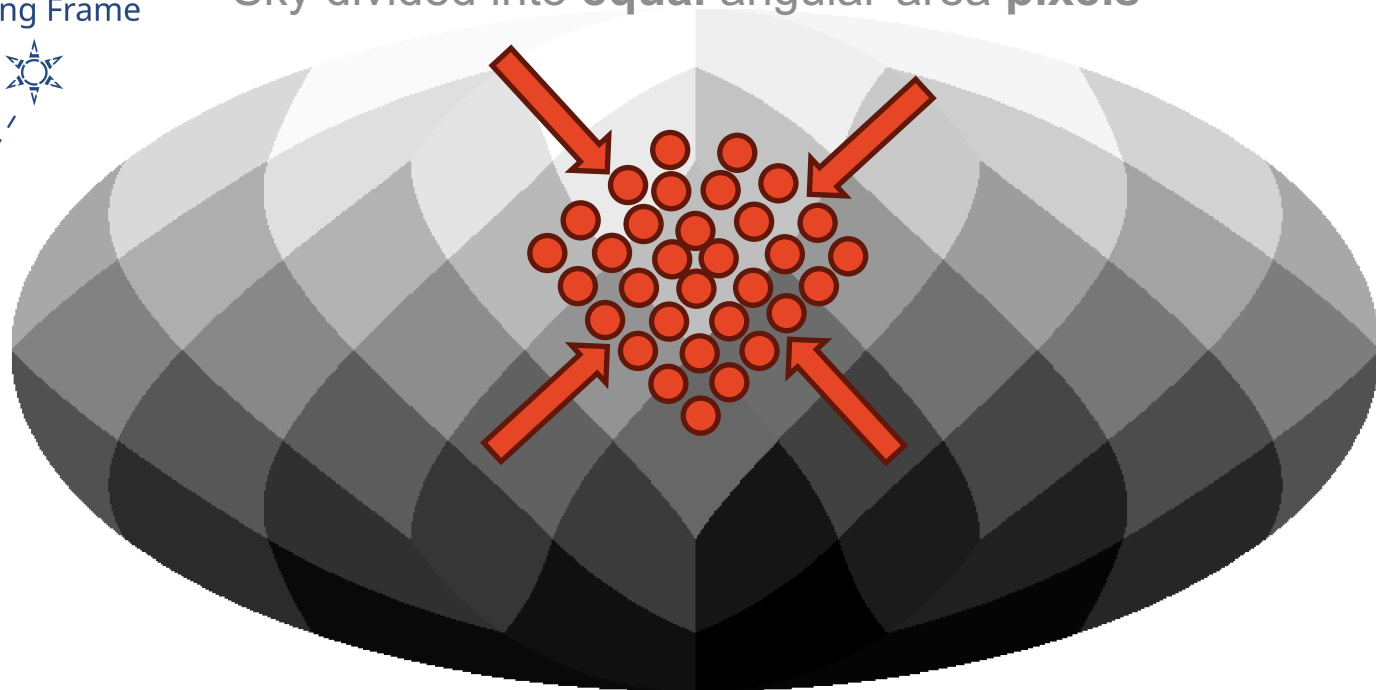


Image credit: Wikipedia

# Relativistic aberration



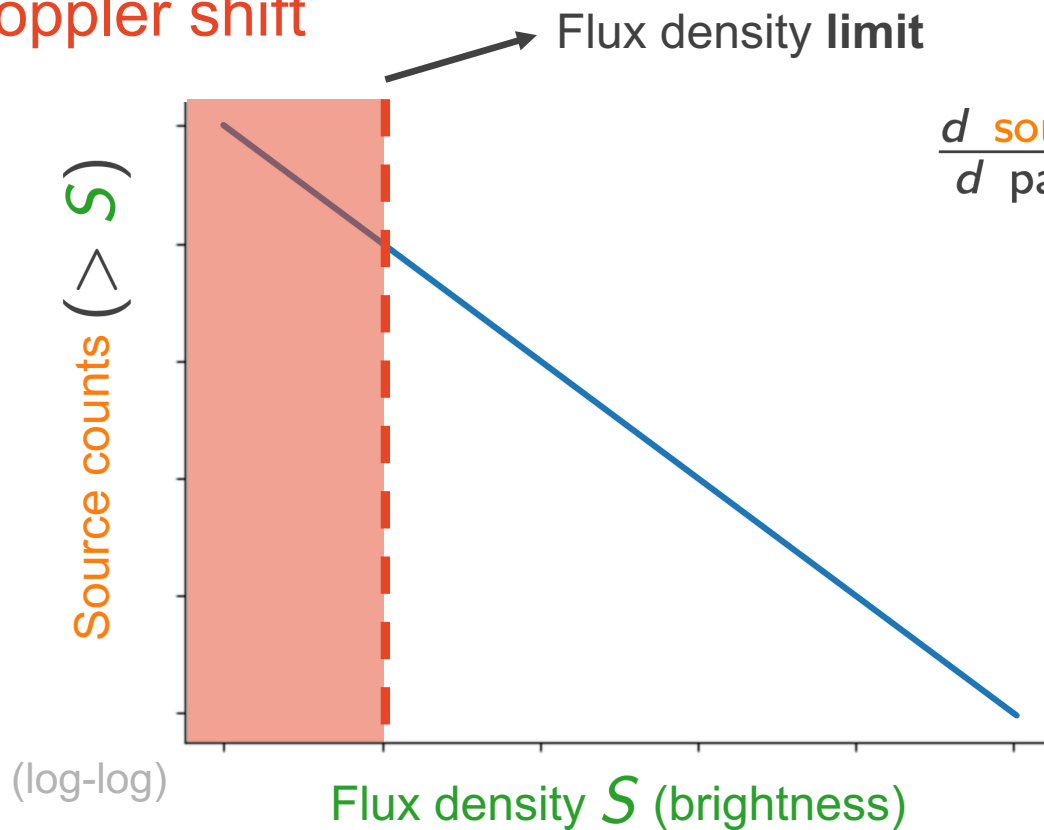
Sky divided into **equal** angular-area pixels



Source **density increases** due to motion

Image credit: Wikipedia

## Doppler shift



$$\frac{d \text{ source counts}}{d \text{ patch of sky}} (> S) \propto S^{-x}$$

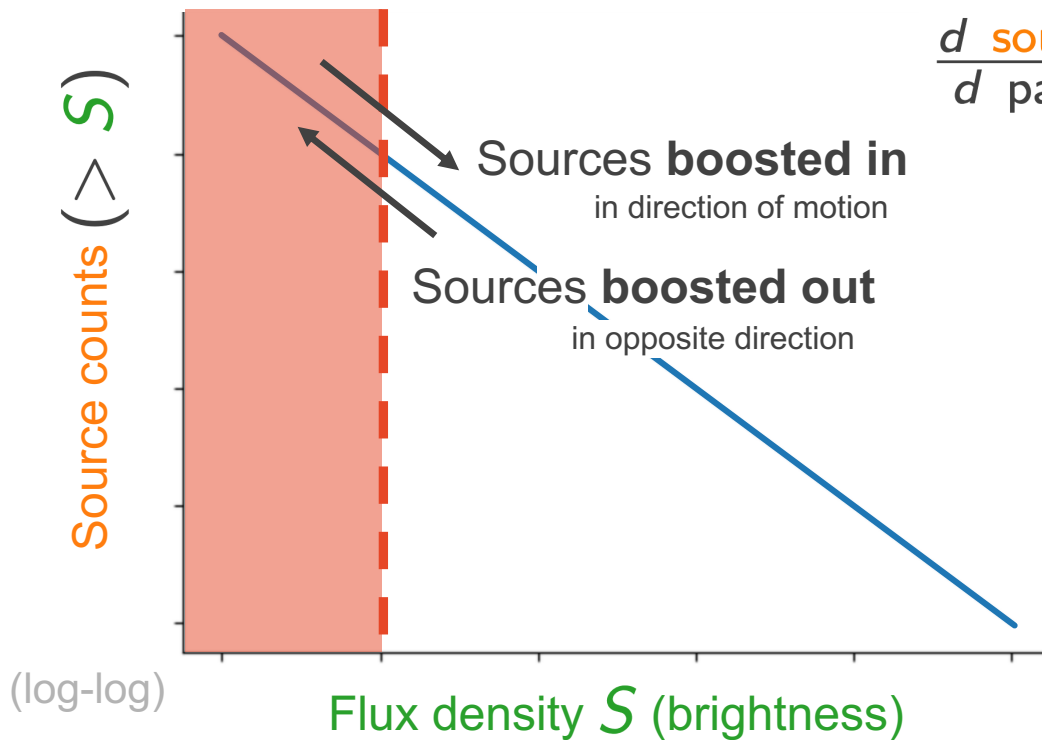
$$S \propto f^{-\alpha}$$

Flux density changes  
with frequency  $f$

Ellis G. F. R., Baldwin J. E., 1984, MNRAS, 206, 377



# Doppler shift



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## The dipole modulation

$$D = (2 + \underbrace{x[1 + \alpha]}_{\text{Doppler}})(v/c)$$

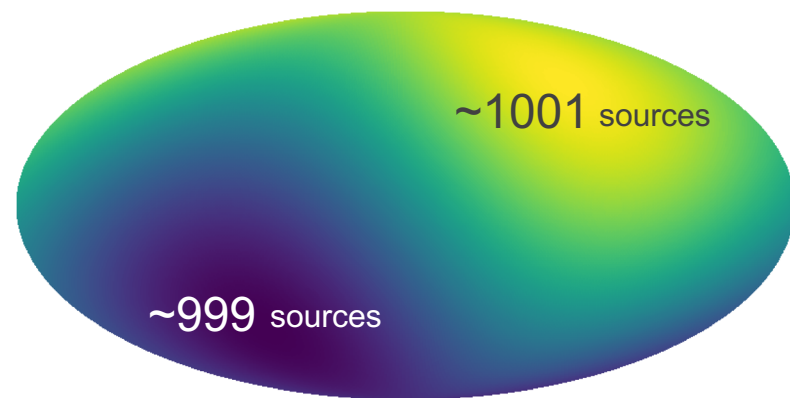
↓  
Aberration

## The dipole modulation

$$D = (2 + \underbrace{x[1 + \alpha]}_{\text{Doppler}})(v/c)$$

↓  
Aberration

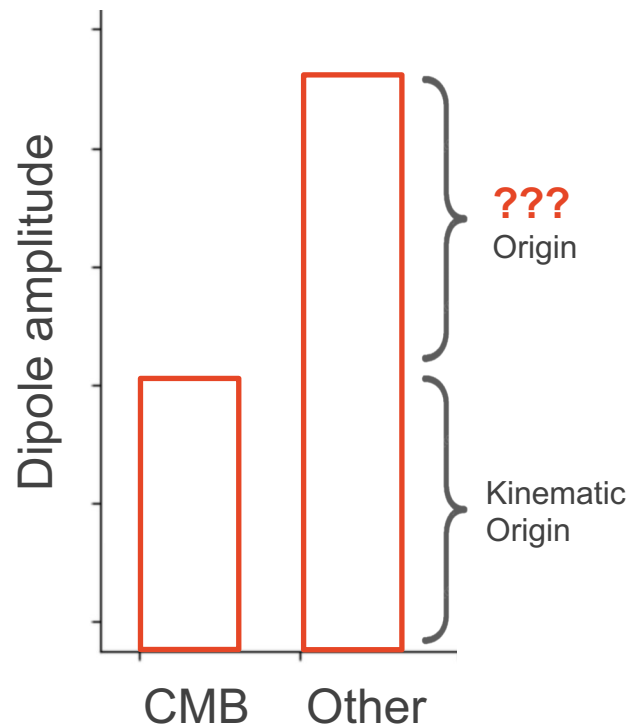
This is a small effect, modulation of  
**around 0.1% over the sky**



Average 1000 sources/pixel  
with **amplitude**  $D \sim 0.001$

## Recap: How does this relate to cosmology?

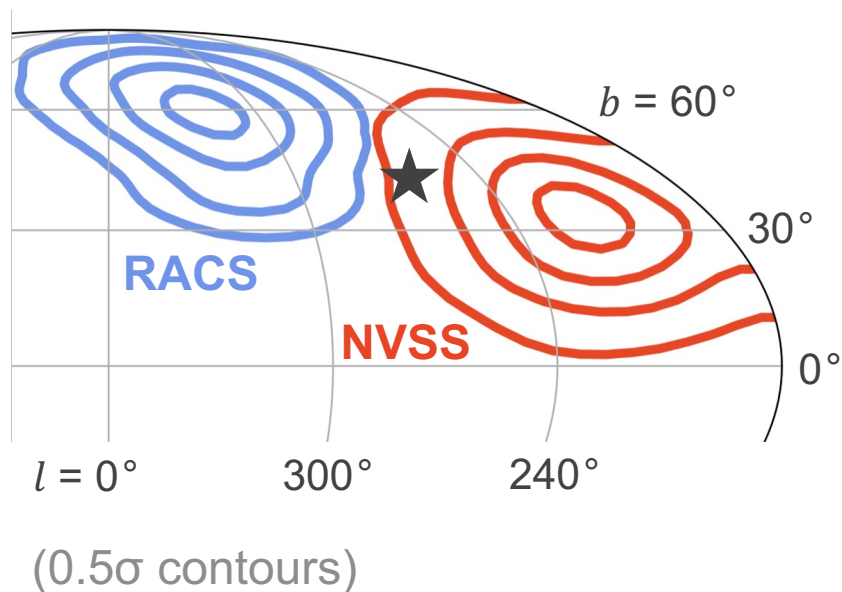
- Can **measure** dipole in survey (e.g. galaxies)
- If dipole **disagrees** with CMB, this indicates an **anisotropy**
- Anisotropy violates the **Cosmological Principle**



# What dipoles have been found?

Dipole **directions**:

- Generally **agree** with the CMB



Dipole **amplitudes**:

- Consistently **exceed** the CMB amplitude

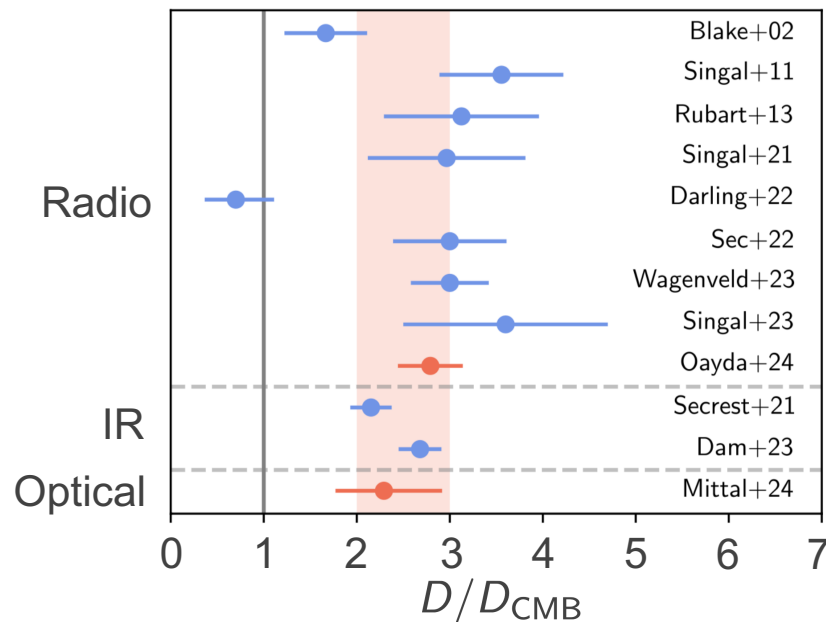
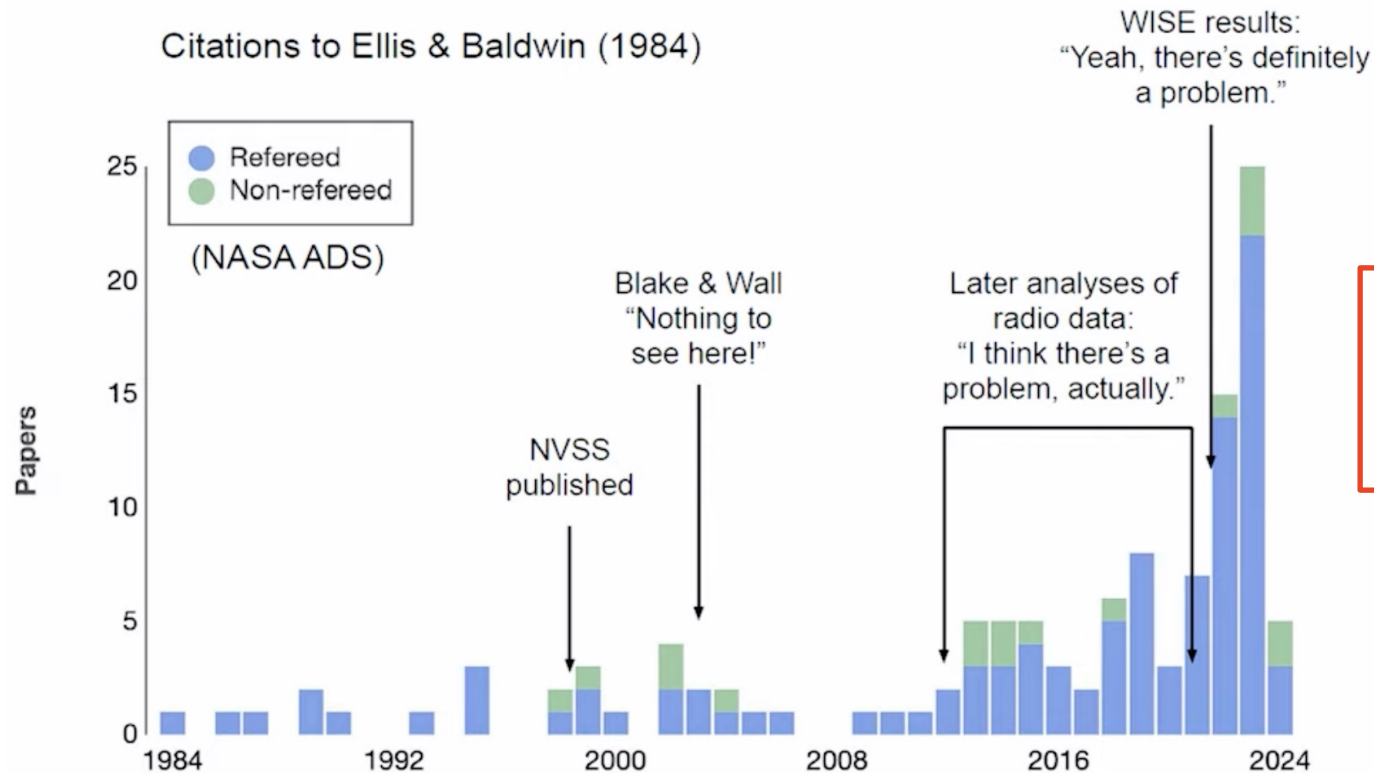


Image credit: Oayda et al. (2025; in prep)

# Houston, we have a problem



Poses a **serious challenge** to the CP and the **standard model**.

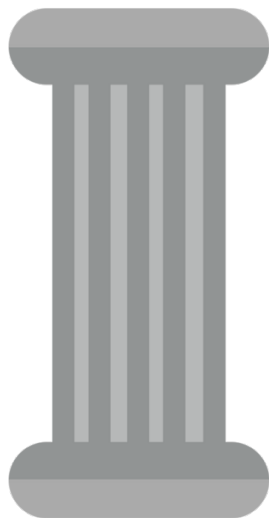
Measurements now **disagree** with the CMB at  **$>5\sigma$** .

This is **comparable** to the **Hubble** tension.

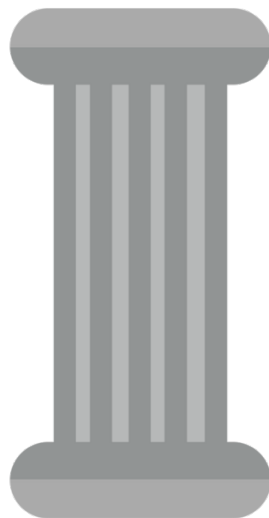
Image credit: Nathan Secrest

# Three pillars of tension

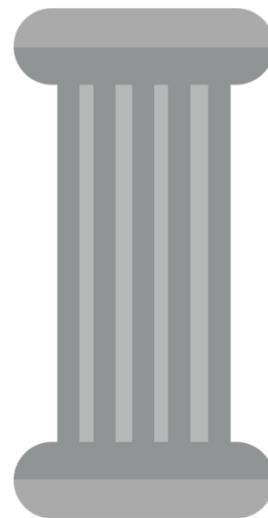
Tension between  
**parameters**



Tension between  
**models**



Tension between  
**datasets**



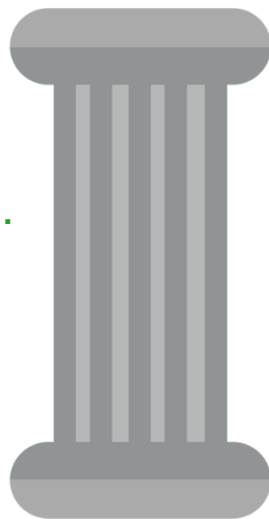
## Three pillars of tension

We do this **here**.

Tension between  
**parameters**

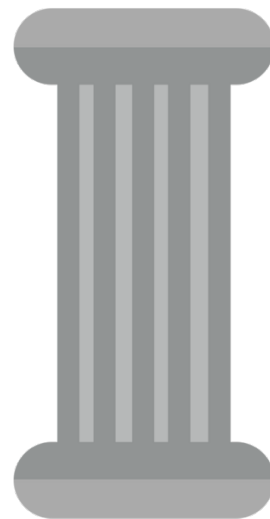


Tension between  
**models**



Already done.

Tension between  
**datasets**





## Tension between datasets: $R$ -statistic

- **Bayesian evidence**  $\mathcal{Z}_D$  is the **probability** of observing the data  $D$

**Concordance:**

$$R \gg 1$$

$$\mathcal{Z}_{AB} \gg \mathcal{Z}_A \mathcal{Z}_B$$

Combining datasets  
**increases** confidence

$$R = \frac{\mathcal{Z}_{AB}}{\mathcal{Z}_A \mathcal{Z}_B}$$

**Tension:**

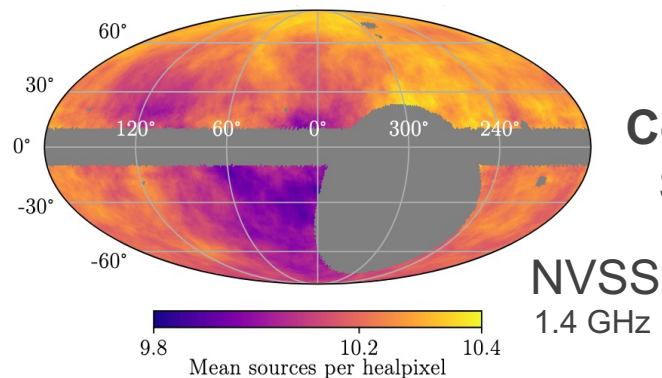
$$R \ll 1$$

$$\mathcal{Z}_{AB} \ll \mathcal{Z}_A \mathcal{Z}_B$$

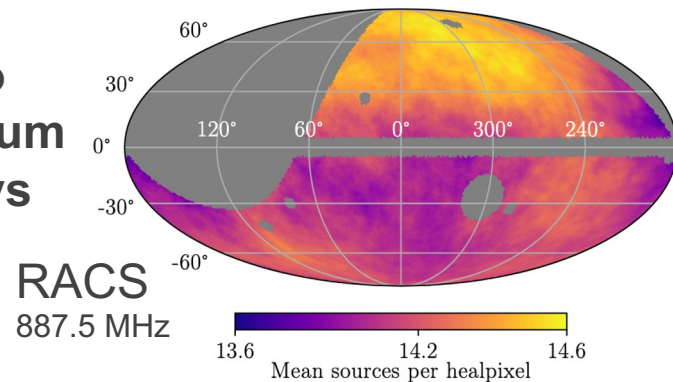
Combining datasets  
**decreases** confidence

Can convert  $R$  to  $\sigma$ -level tension

# Surveys of interest...

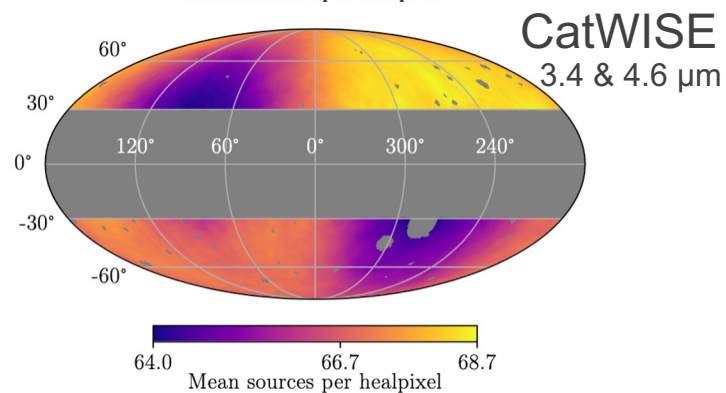
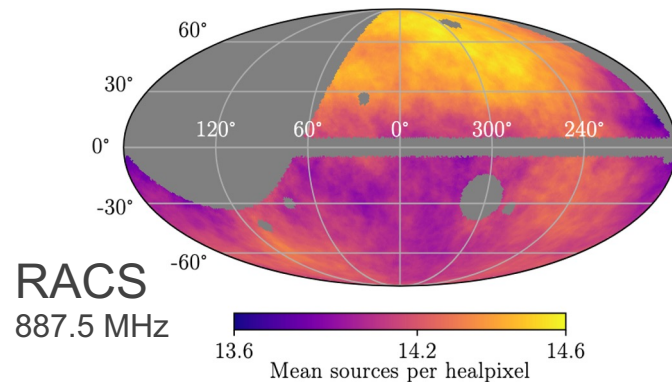
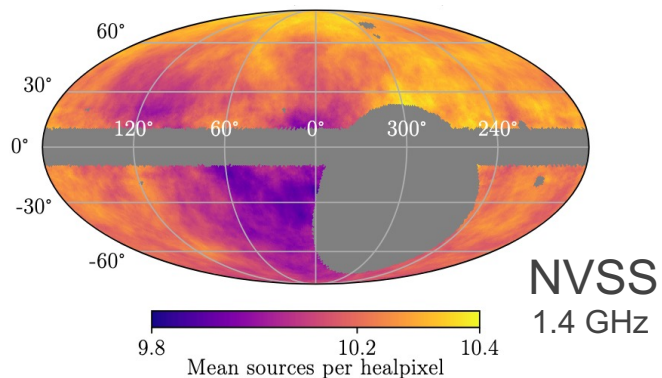


## Radio Continuum Surveys

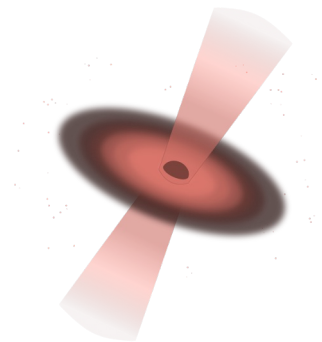


NB: Radio/IR projections are box-car smoothed within 1sr

# Surveys of interest...

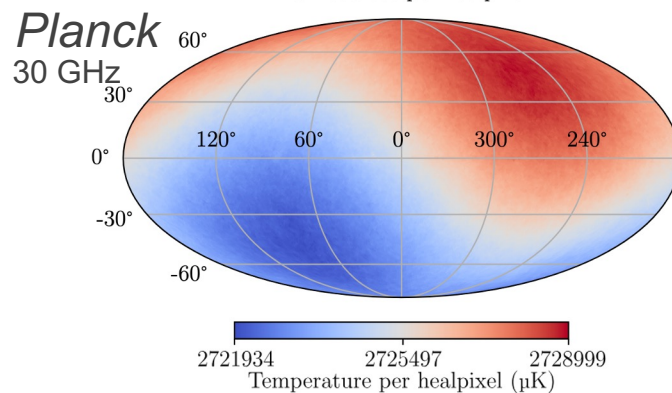
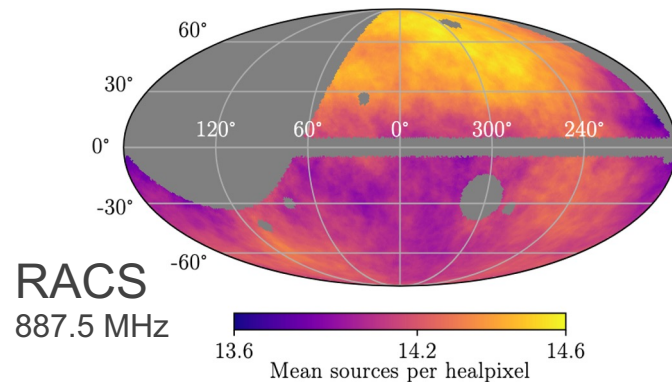
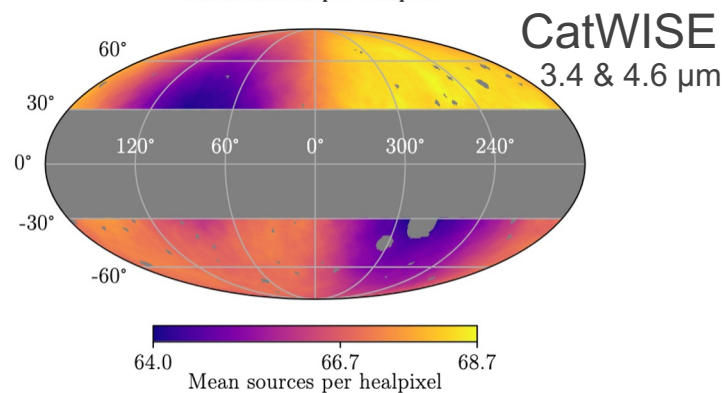
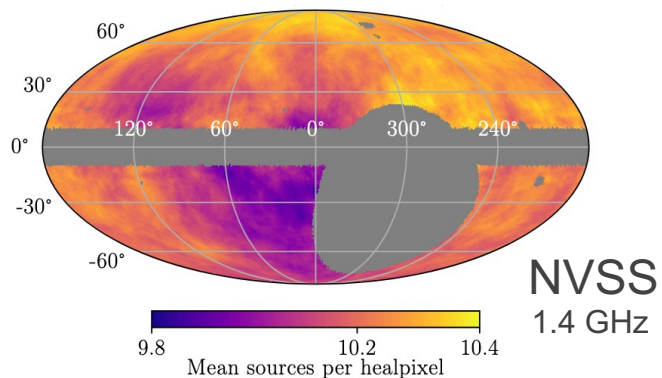


**Infrared-selected  
Quasars**



NB: Radio/IR projections are box-car smoothed within 1sr

# Surveys of interest...



NB: Radio/IR projections are box-car smoothed within 1sr

# The Tension.

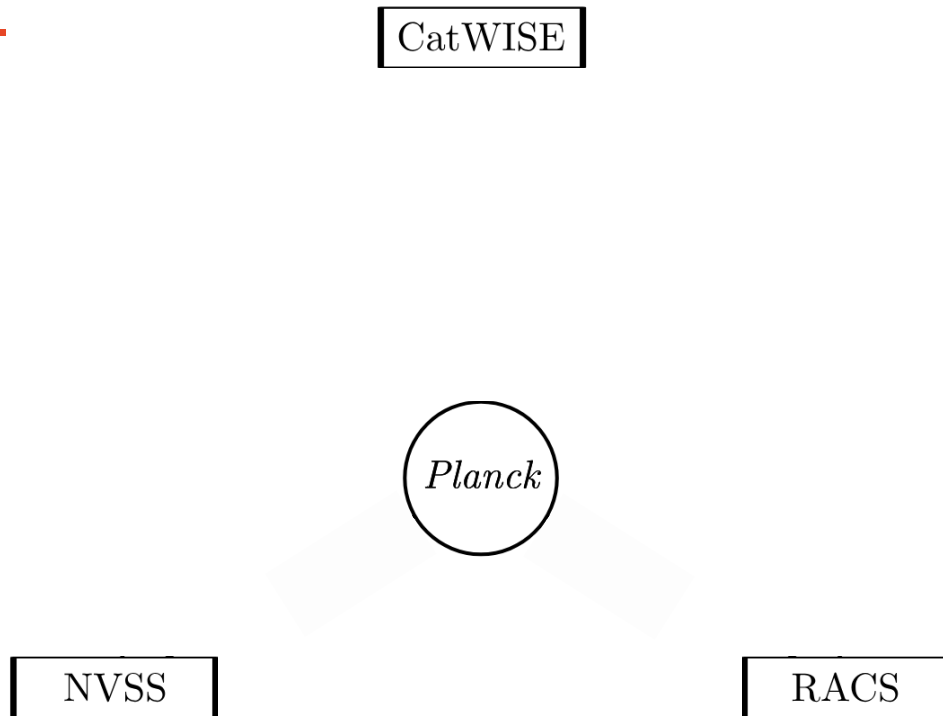


Image credit: Land-Strykowski et al. (2025; submitted)

# The Tension.

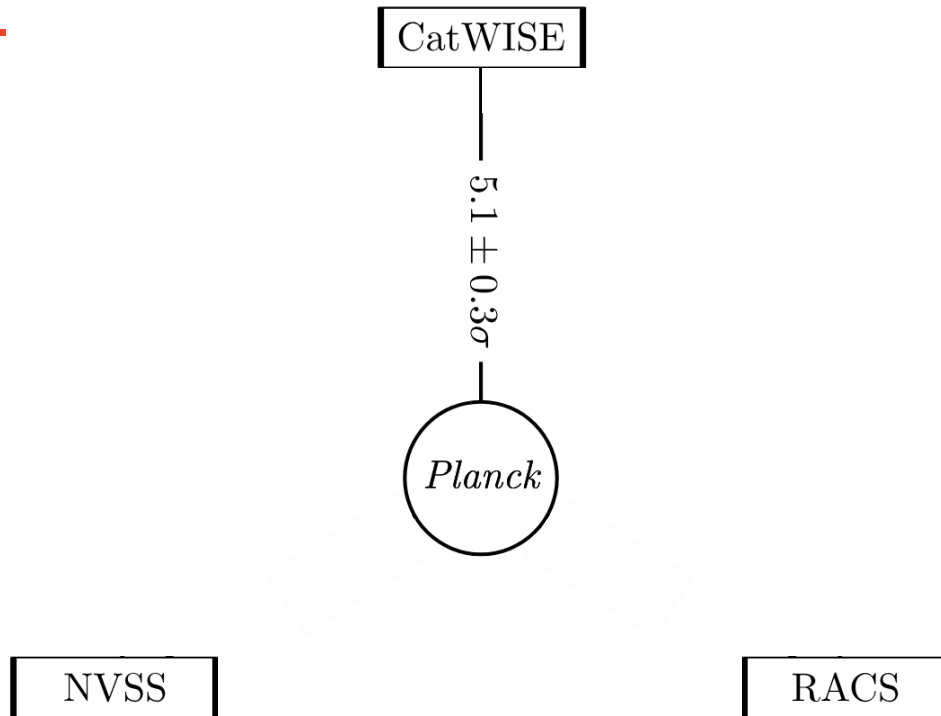


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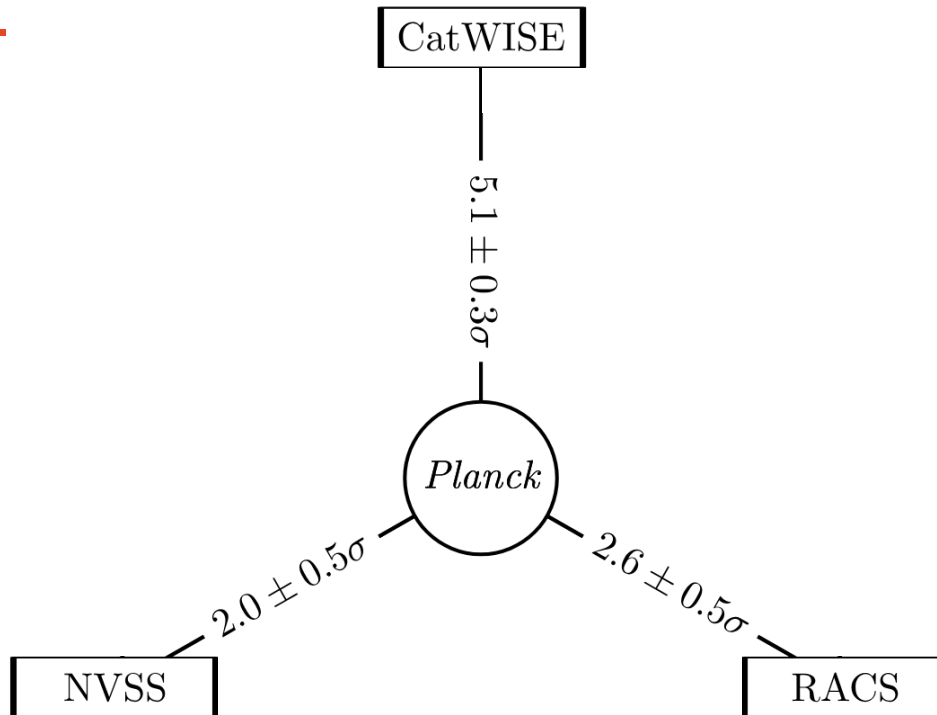


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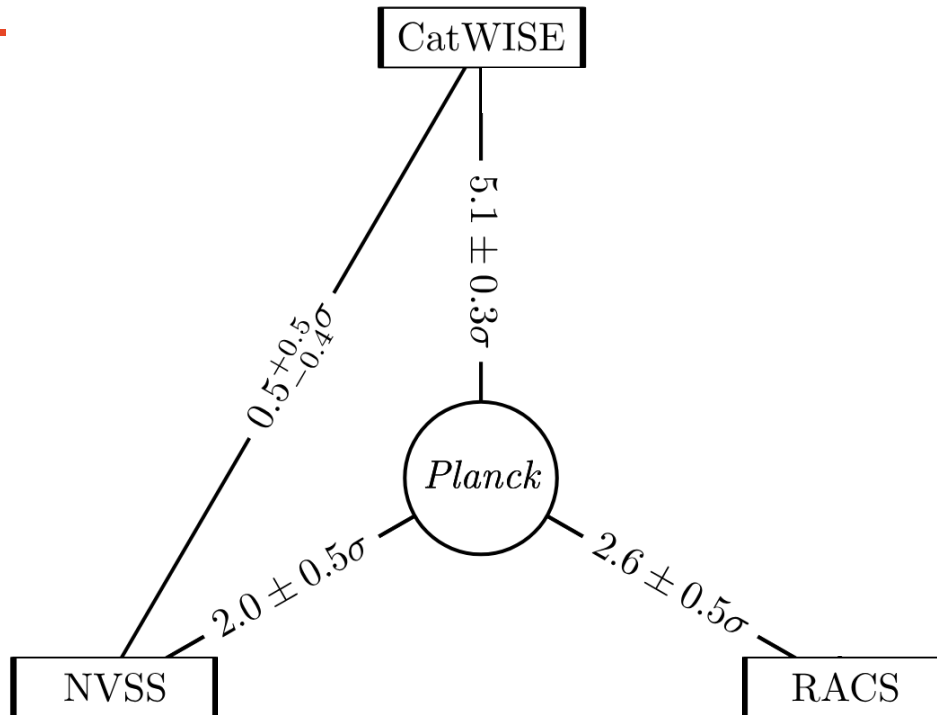


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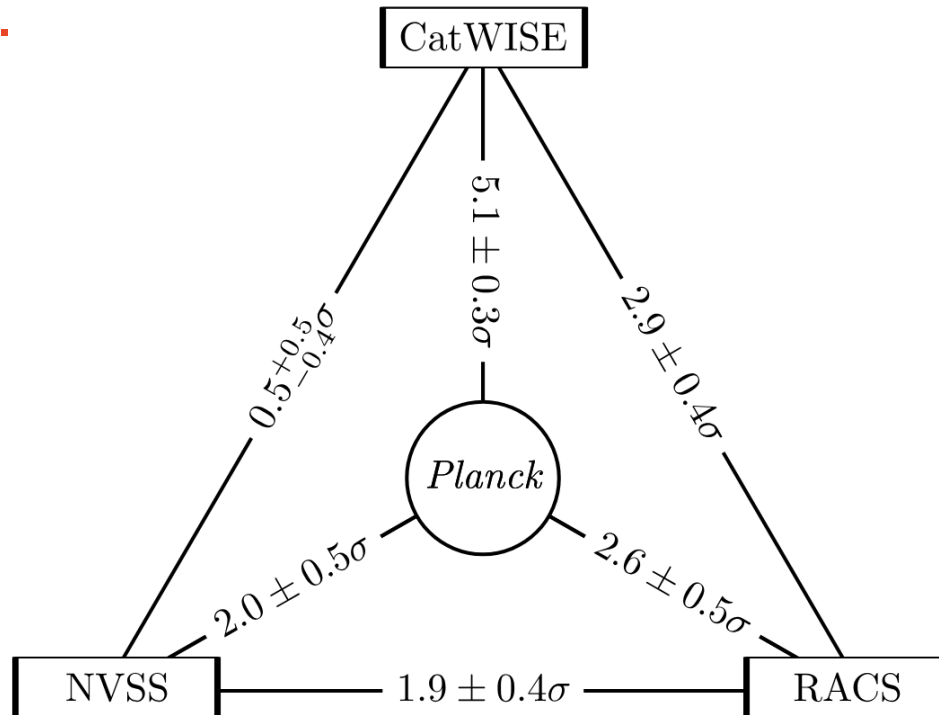


Image credit: Land-Strykowski et al. (2025; submitted)

## Further investigation: RACS

Inference is **sensitive** to:

- Ionospheric temperature fluctuations and **flux-scaling** compensation
- Choice of **resolution** / convolution

Modelling/understanding these is the **future**.

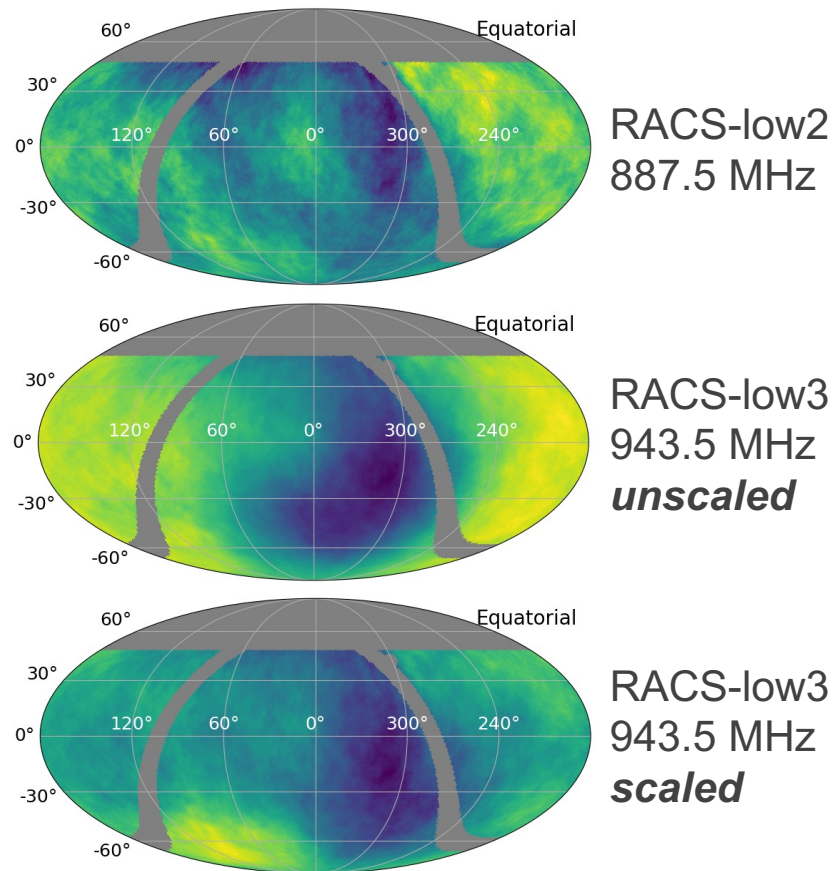
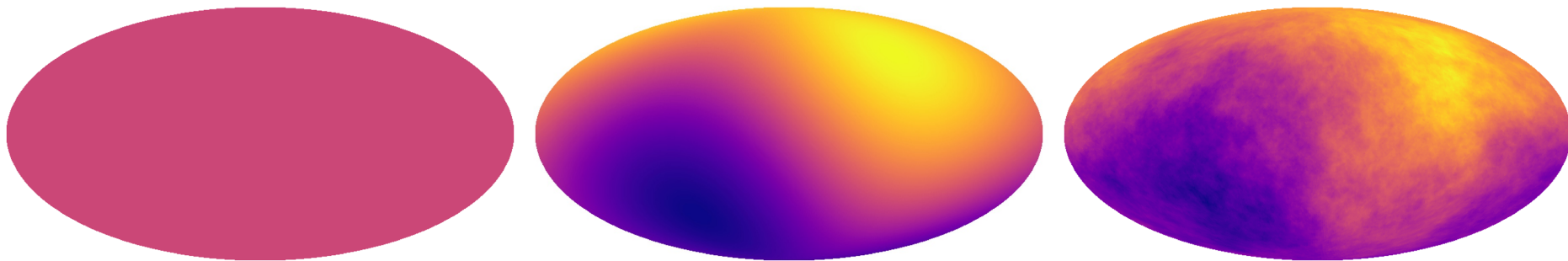


Image credit: Oliver Oayda (2025; private communication)

## Simulated skies: modulation, information



Homogenous sky



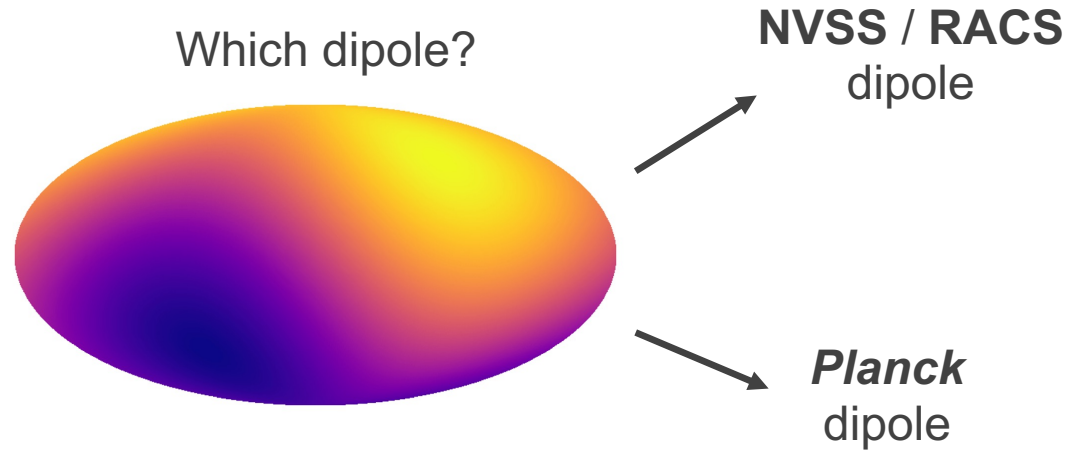
Dipole modulation



Poisson draw

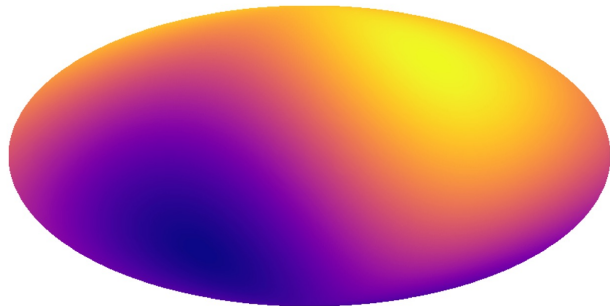
Repeat for **50,000** to **50,000,000** sources

# Simulated skies: different dipoles



# Simulated skies: different dipoles

Which dipole?



**NVSS / RACS**  
dipole

vs

**Planck**  
data

**“Radio continuum dipole”**

**Planck**  
dipole

vs

**Planck**  
data

**“CMB dipole”**

# Simulated skies: tension results

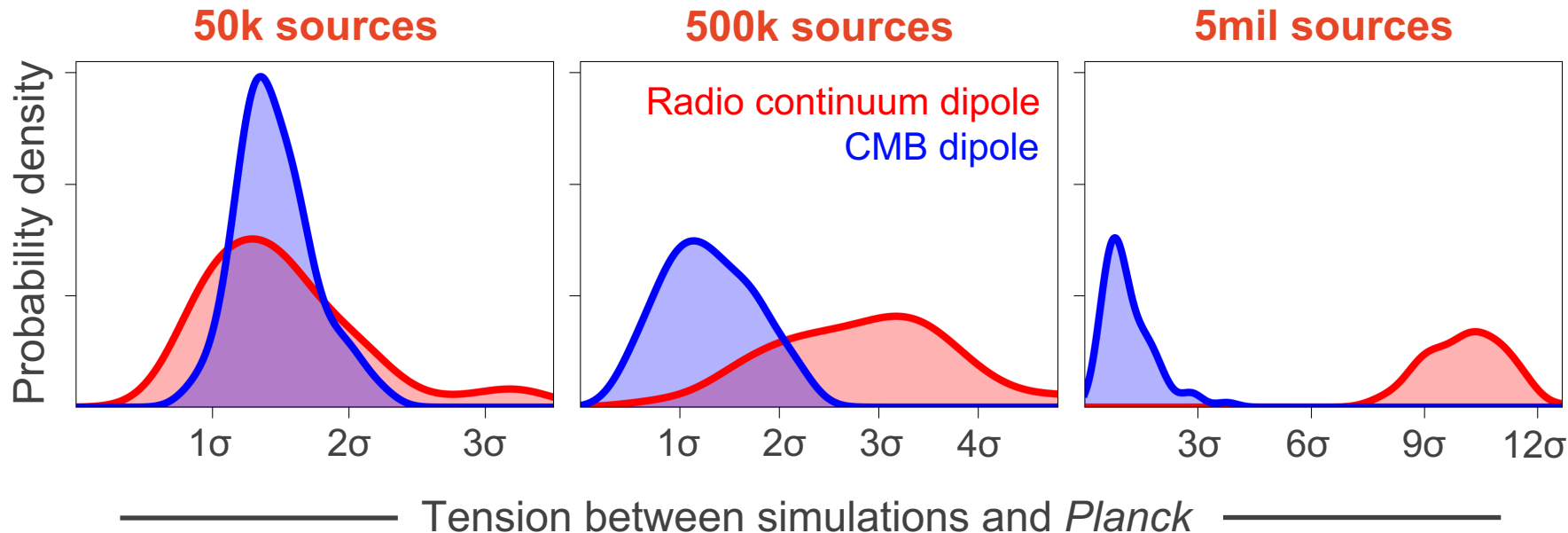


Image credit: Land-Strykowski et al. (2025; submitted)

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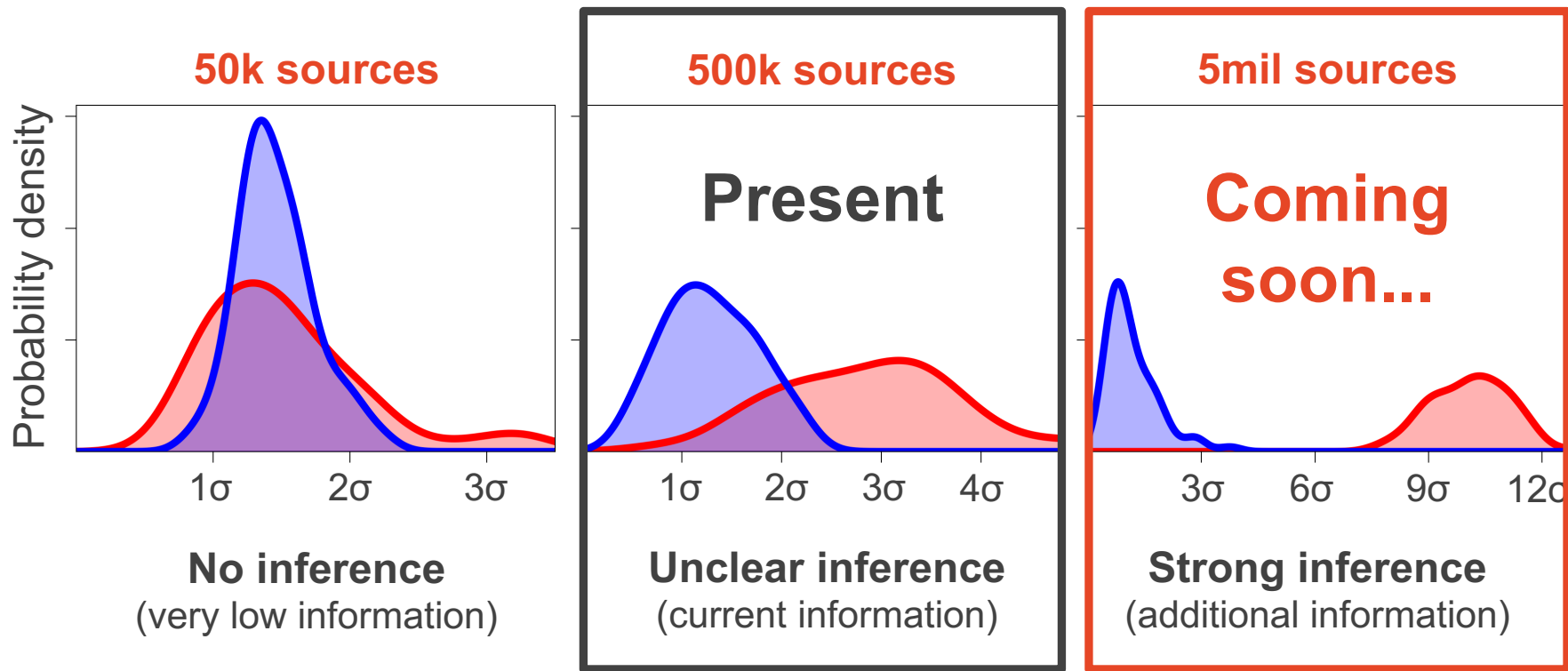


Image credit: Land-Strykowski et al. (2025; submitted)

# How many sources do we need?

- **NVSS** dipole  $\longrightarrow$   **$5\sigma$**  tension with  **$\sim 1,300,000$**  sources
- **RACS** dipole  $\longrightarrow$   **$5\sigma$**  tension with  **$\sim 1,500,000$**  sources

Only need  **$O(10^6)$**  more sources!!

SKA will observe  **$5 \times 10^6$**  and  **$9 \times 10^8$**  galaxies

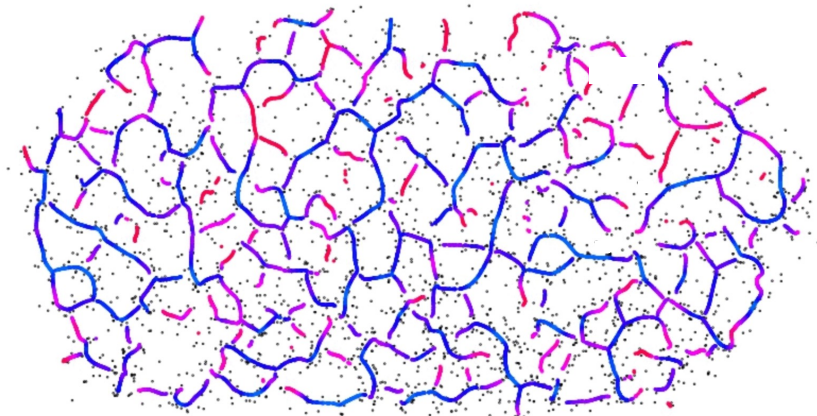
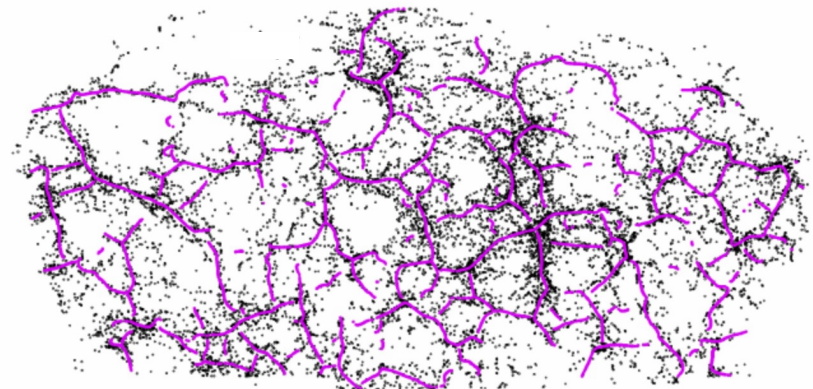


We are on the cusp of untangling the anomaly...



## Change is on the horizon...

- Why homogeneity / isotropy?
- **Inhomogeneous** cosmologies...
  - Dark matter super-void (1–3.4 Gpc)
- **Anisotropic** cosmologies...
  - Tilted cosmology (not Friedmann)
- Or, perhaps, something entirely new...



# Conclusions

- The **Cosmic Dipole tension** appears to be one of the **most significant** tensions.

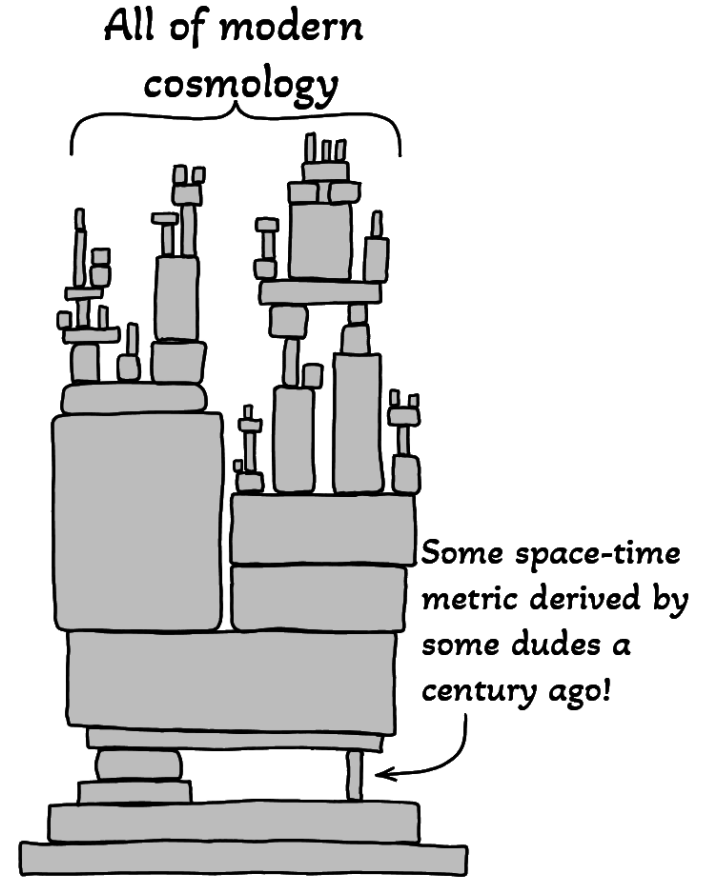


Image credit: Geraint F. Lewis & xkcd

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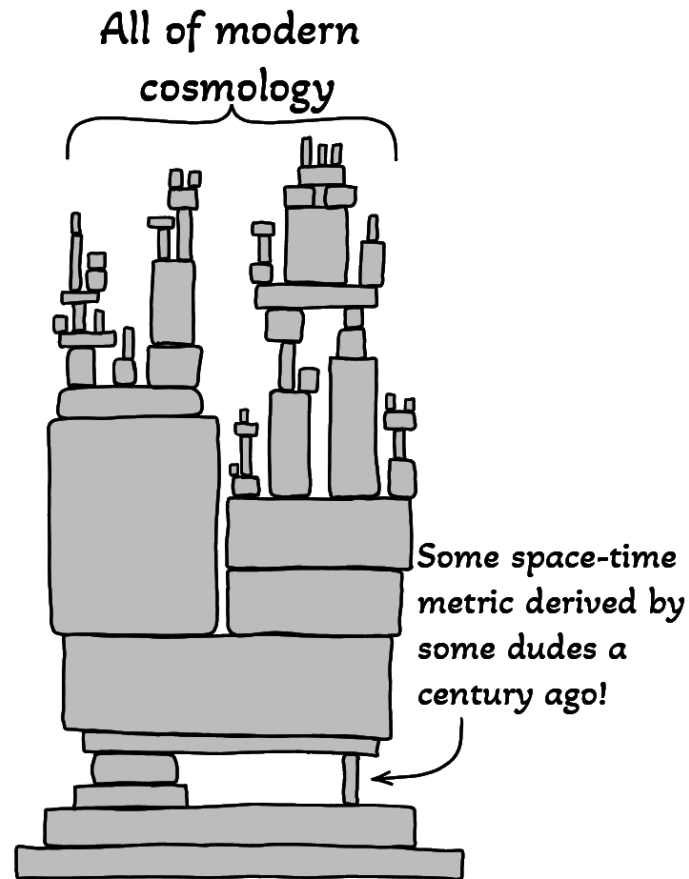


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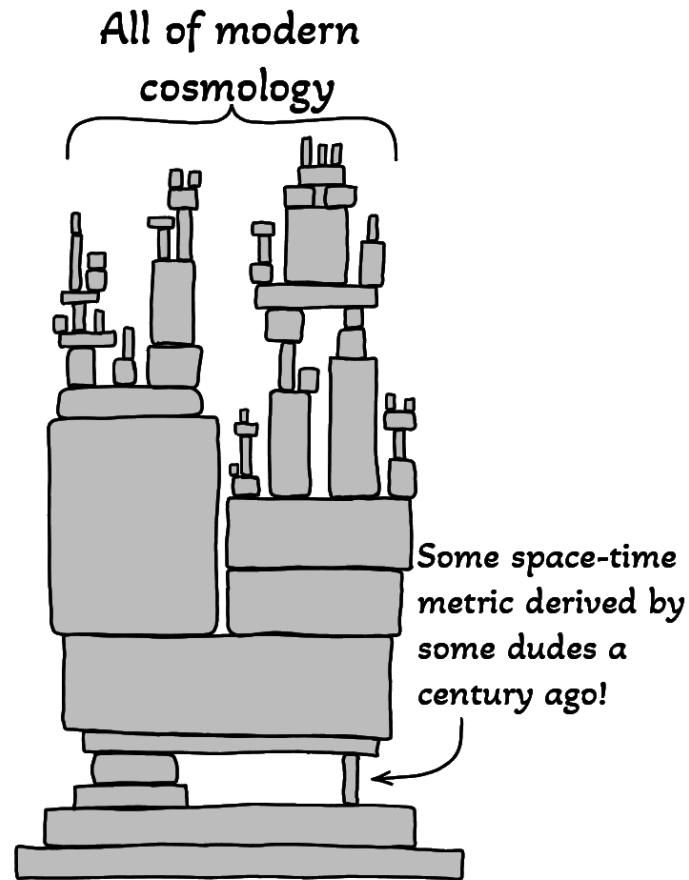


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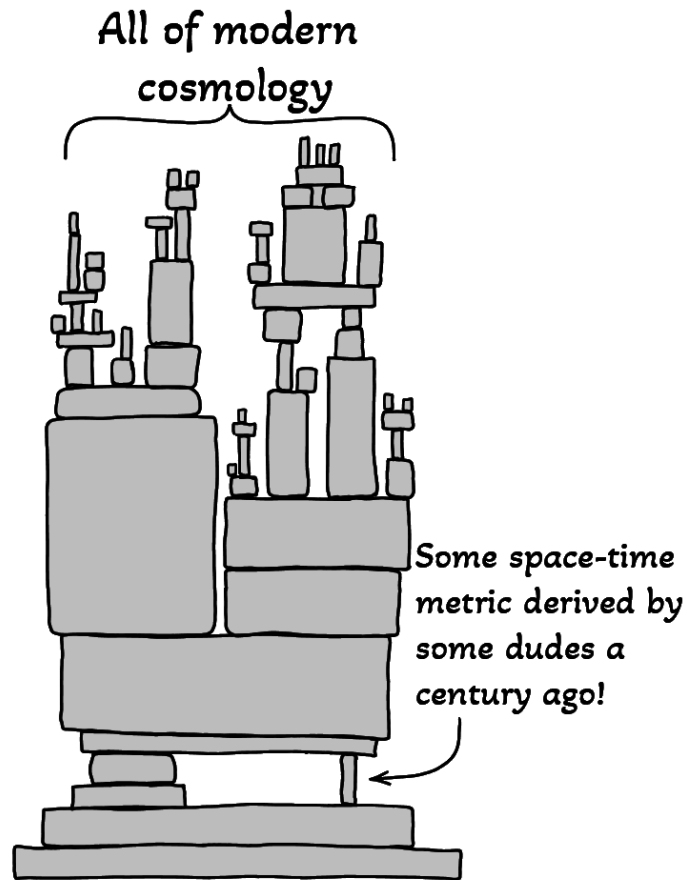


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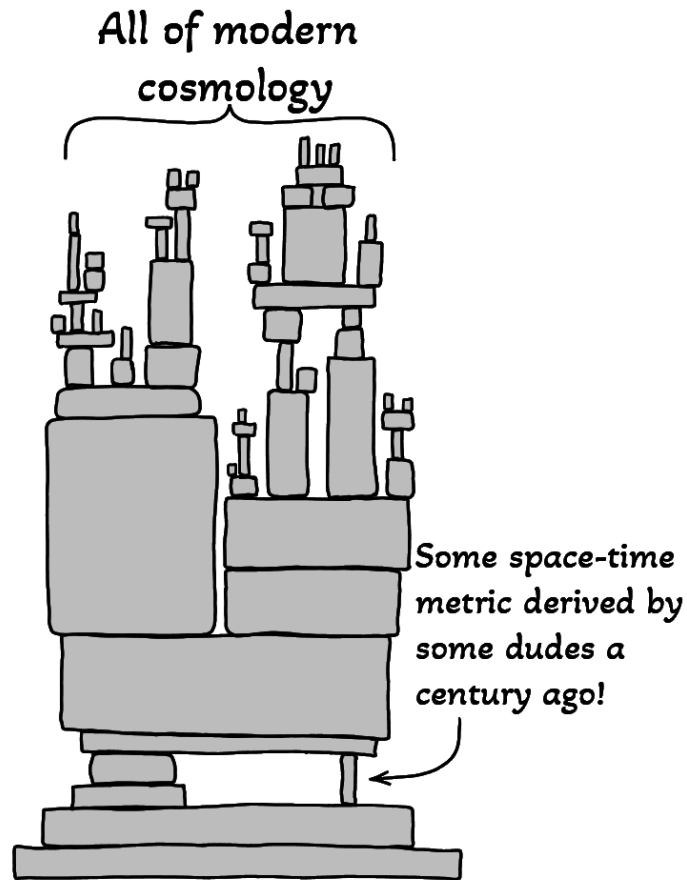


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Thank you.  
Questions?

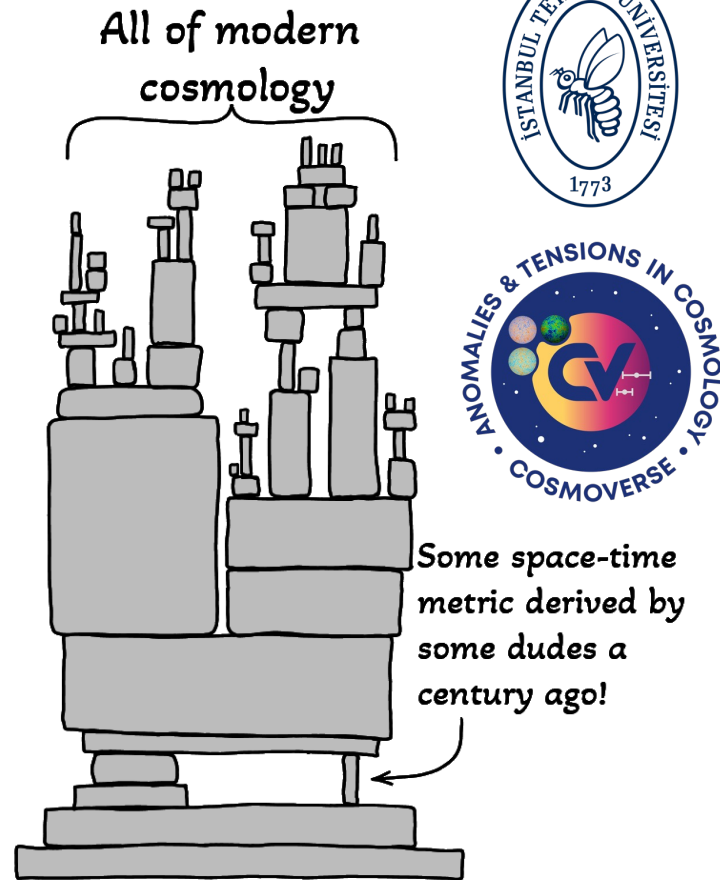


Image credit: Geraint F. Lewis & xkcd